



Current status of Xenoscope, a full-scale vertical demonstrator for the DARWIN observatory

Nagoya Workshop on Technology and Instrumentation in Future Liquid Noble Gas Detectors, 2024.

Jose Cuenca-García, on behalf of the Xenoscope team.

Two demonstrators for DARWIN

Pancake (Freiburg)

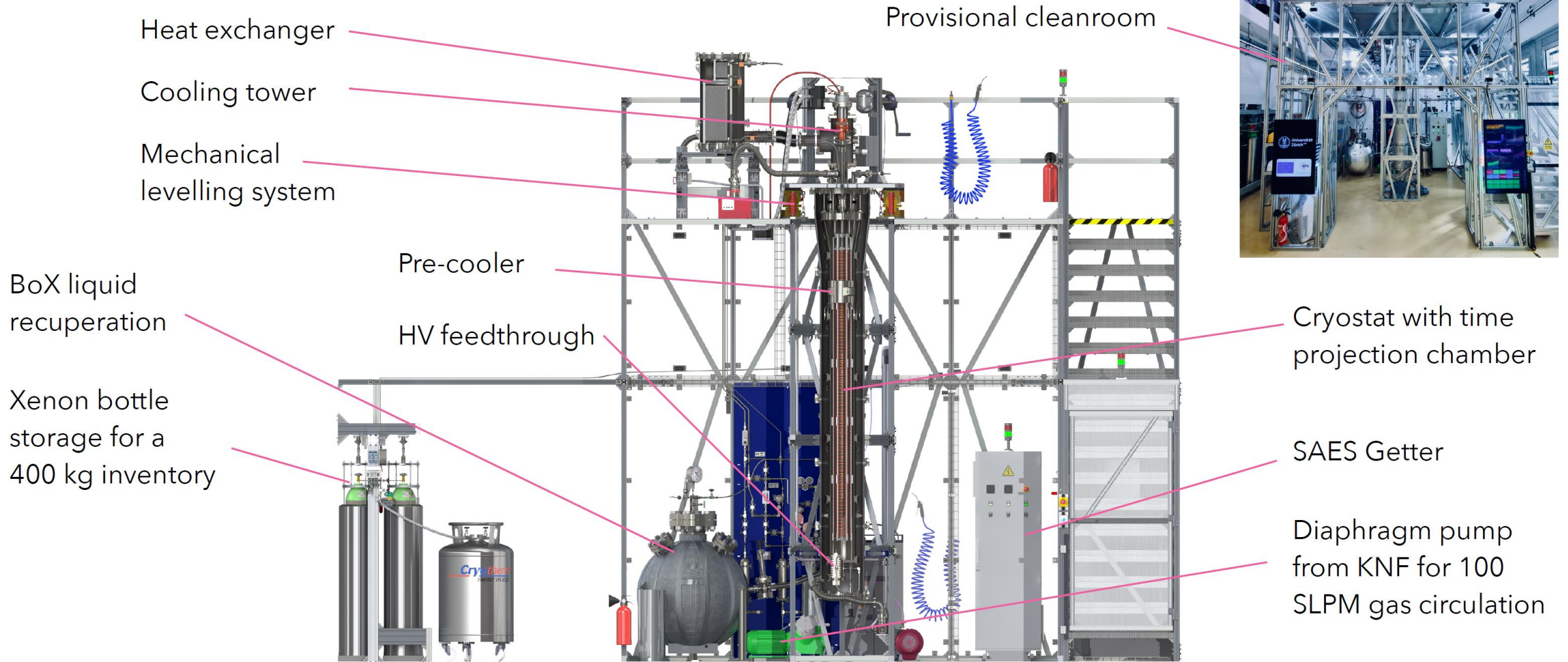


Xenoscope (Zurich)



Xenoscope

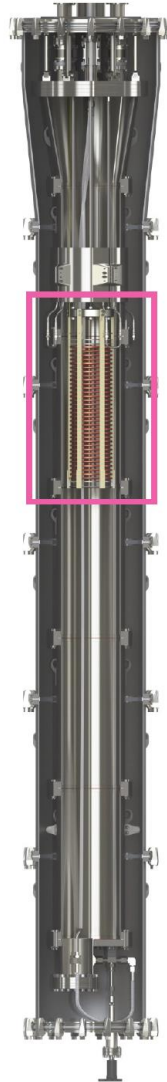
JINST 16 P08052 (2021)



The plan

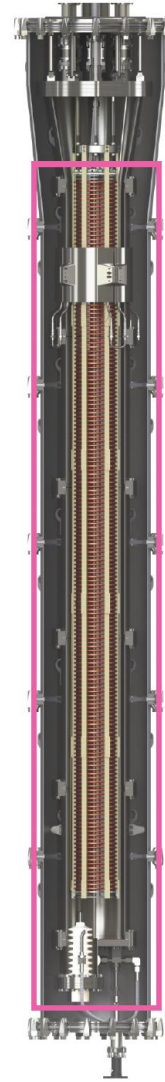
Phase 1: purity monitor

- Stable operation for 3 months
- Test cryostat systems (pre-cooler)
- Test of liquid recuperation to BoX
- Model electron lifetime as a function of purification speed/conditions
- Model electron transport properties for different electric fields



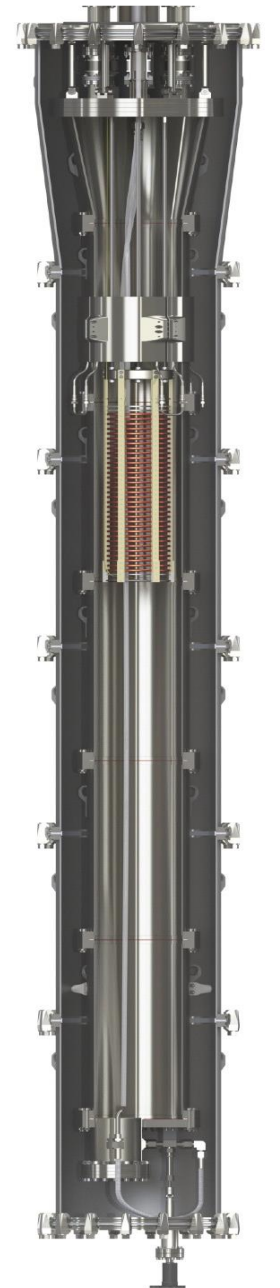
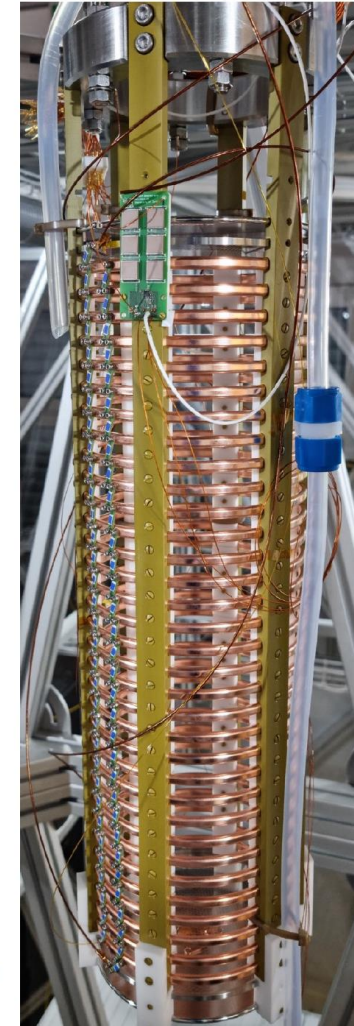
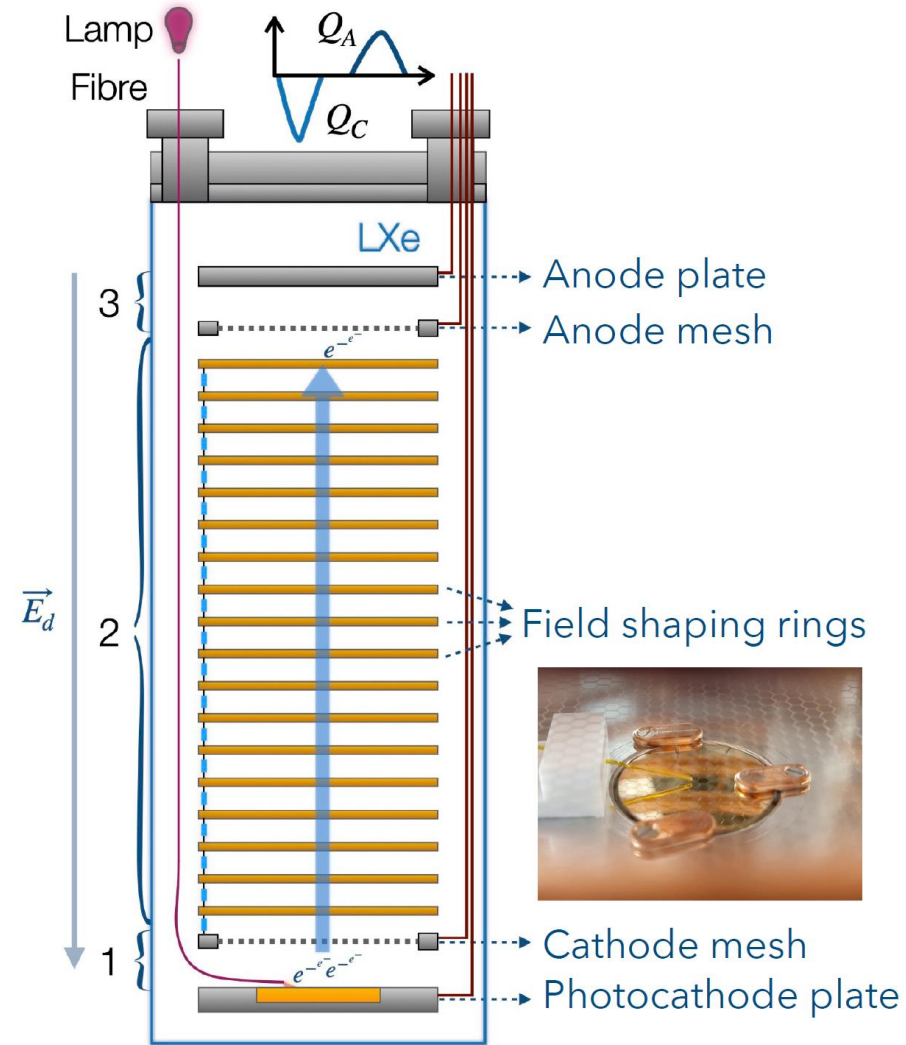
Phase 2: 2.6-m Time Projection Chamber

- High-voltage up to 50 kV
- SiPM array on top to detect S2 light
- Liquid level control
- Additional calibration sources (muons, external sources) to further improve our models



Phase 1: purity monitor

- Drift length of ≈ 50 cm
- Gold-coated quartz photocathode produced in house
- A xenon flash lamp (190 – 2000 nm) was used to illuminate the photocathode
- The FSR provide an homogeneous drift field of 25 - 75 V/cm
- No light: we measured the induced currents



Results on: [Eur. Phys. J. C \(2023\) 83: 717](#)

Jump to next step

Several upgrades were made from the purity monitor

Field cage



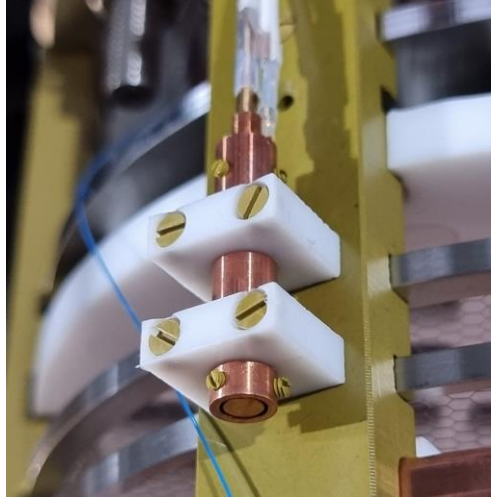
High-voltage



SiPM array



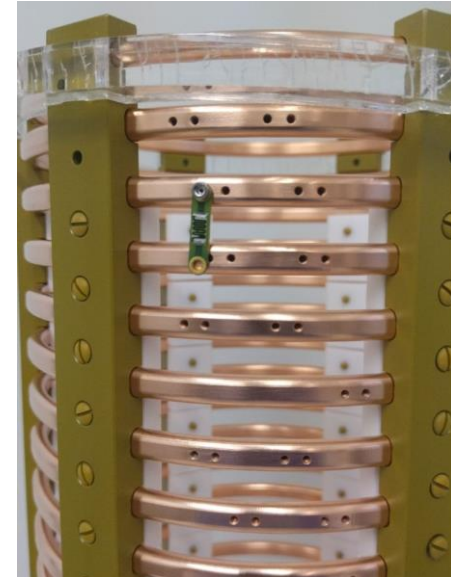
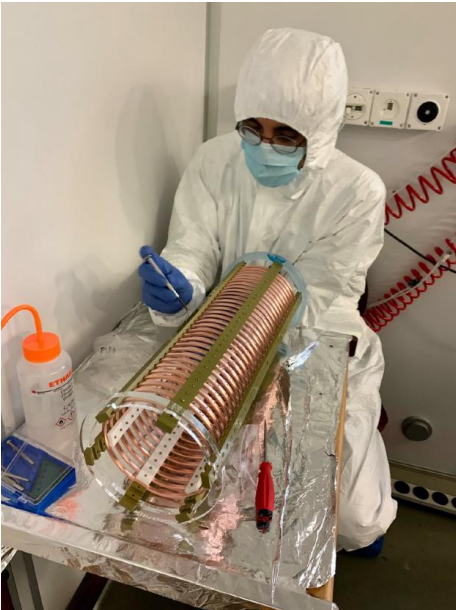
Liquid level control



Upgrade: the field cage



- Divided into 5 independent modules (52 cm each)
- In total, we have 173 copper rings
- Pillars made of Torlon and PTFE inner connectors
- Two parallel 1 G Ω resistor chains (200 nA at 20kV)

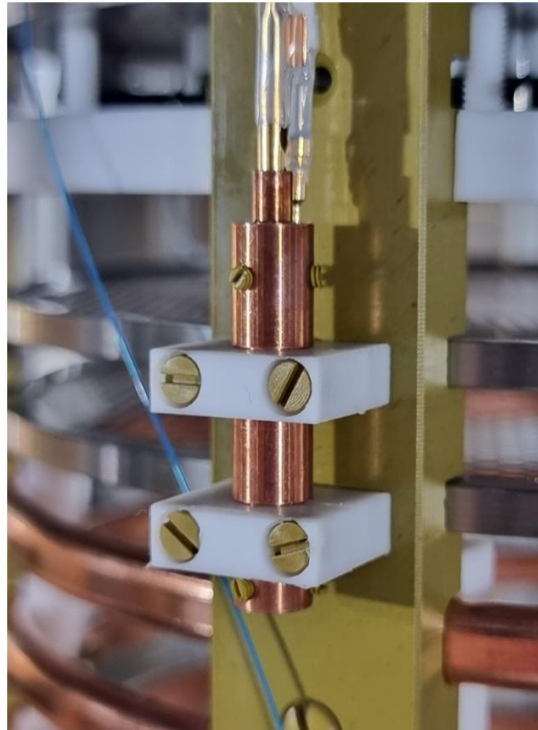


Upgrade: liquid level control

Long Level Meter (LLM)



Short Level Meter (SLM)



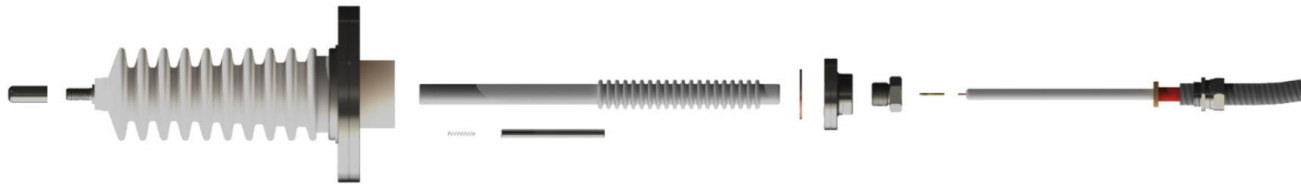
Weir



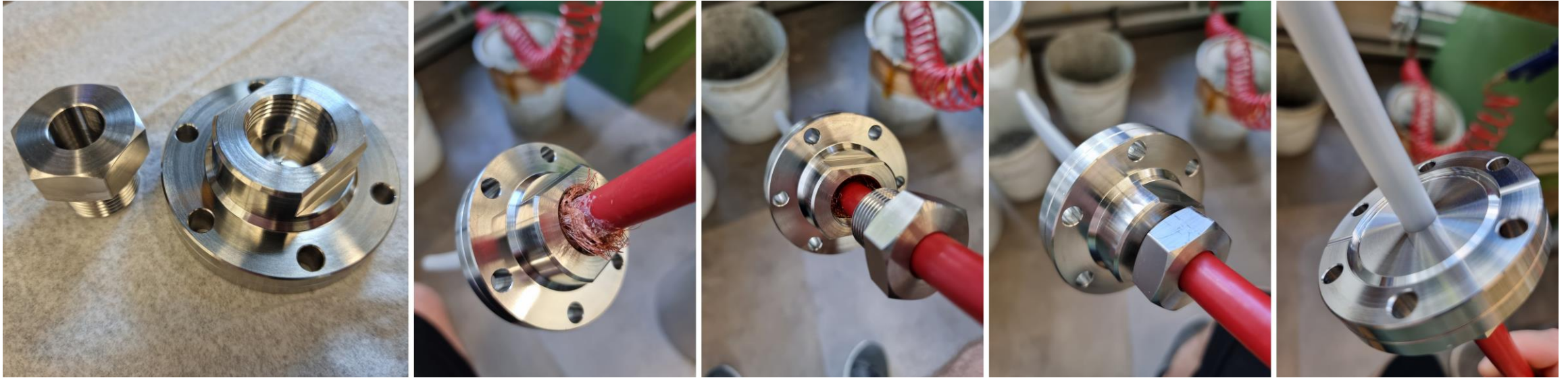
- Level-dependent capacitance of cylindrical level meters
- 3 SLMs + 2-piece segmented LLM
- Active level control with weir on motion feedthrough
- Plan to integrate it in SC readout

Upgrade: high-voltage feedthrough

- HV enters the TPC from the bottom (bending radius ~ 28 cm)
- Commercial Ceramseal (CeramTec) FT, rated 100 kV. Inhouse cryofitted air-to-vacuum FT
- HV rating improved by conditioning, HMWPE insulation elongation, surface treatment
- Entire assembly was tested in vacuum at 120 K and ~ 4 bar overpressure
- The HVFT was successfully operated at 50 kV for over 4 days without sparks



The cryo-fitting



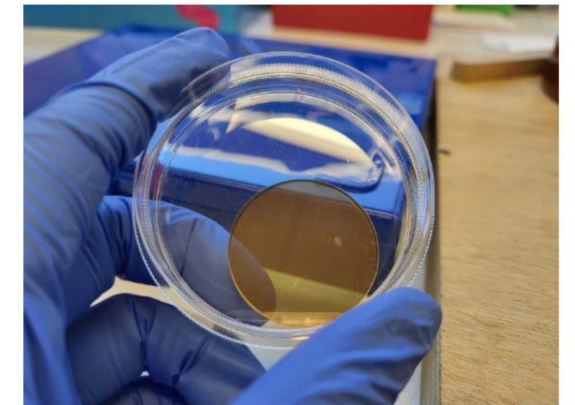
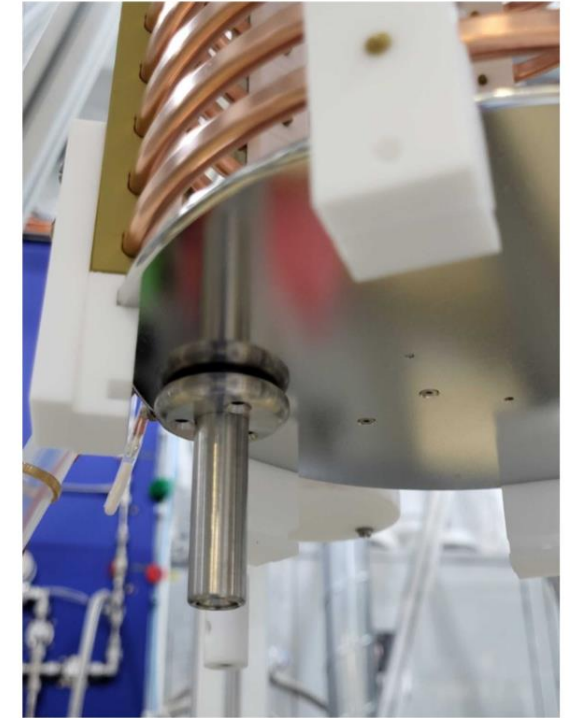
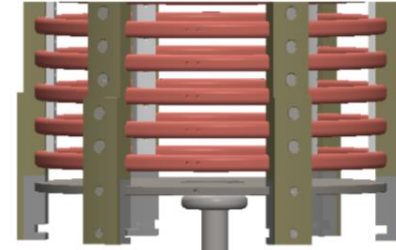
- The cable (Heinzinger HVC100) is stripped of the last three layers
- The white HMWPE insulator (\varnothing 10.7 mm) is cryo-fitted to the custom DN40CF flange
- The cable is slowly dipped in LN2 and cooled down to 77 K
- When the cable warms up to room temperature, it gets cryo-fitted in the flange, making a vacuum tight seal

Photocathode and high-voltage connection

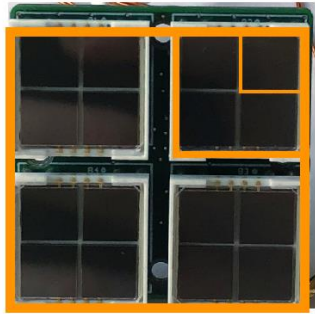
- New photocathode gold coated (50 nm thickness) on a quartz substrate (2 mm thick)
- Xe flash lamp pulse transmitted through optical fibre
- Round-edged HV terminal connected to the Ceramic FT delivers HV to cup & spring system
- Cup & spring system for HV transmission to the cathode



Q150T S automatic sputter coater

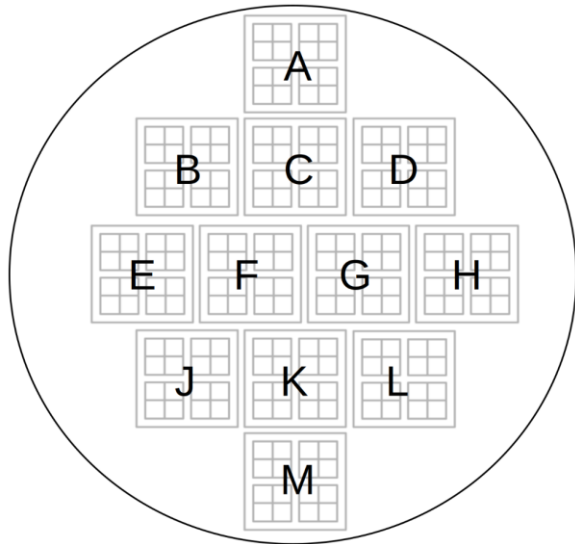
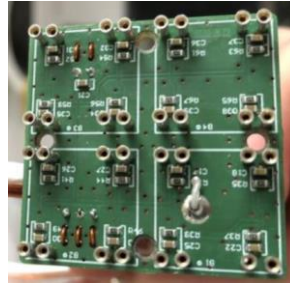


Upgrade: SiPM array

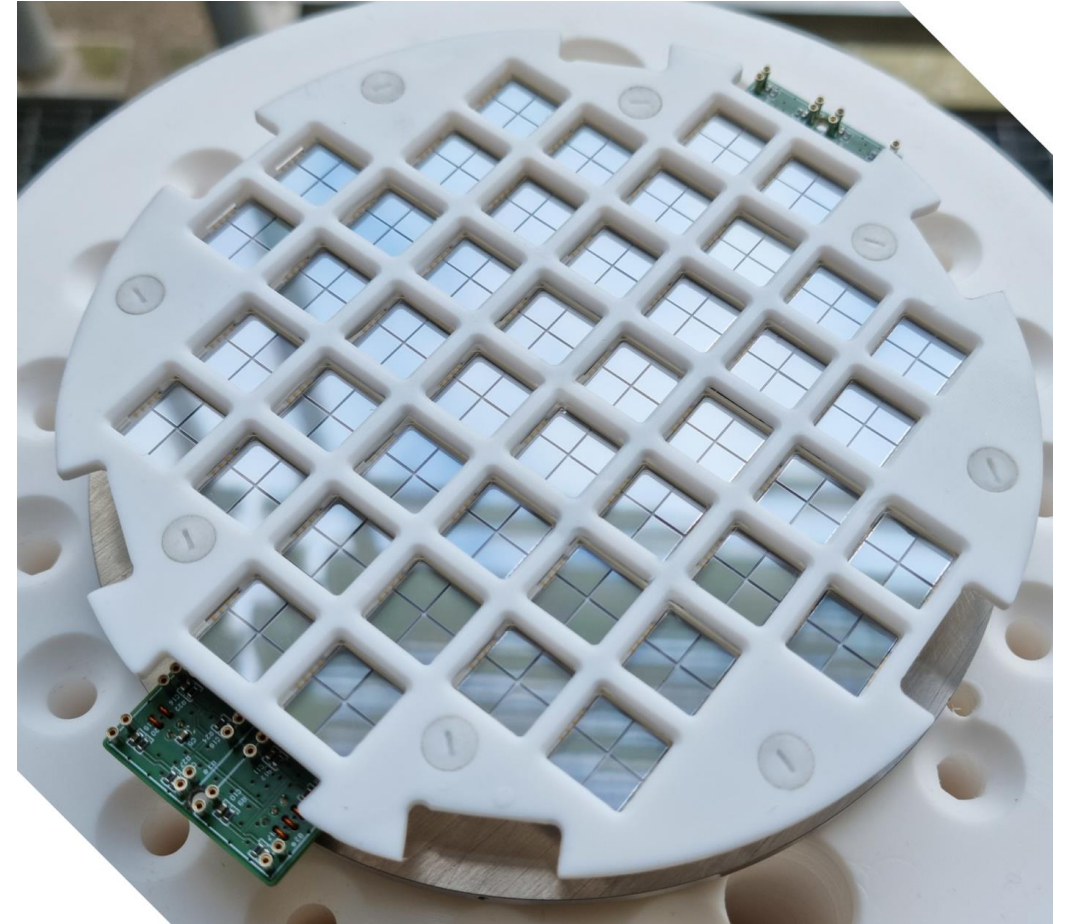


← 34 mm →

Cell = 6 x 6 mm²
Quad = 4 cells
Tile = 4 quads

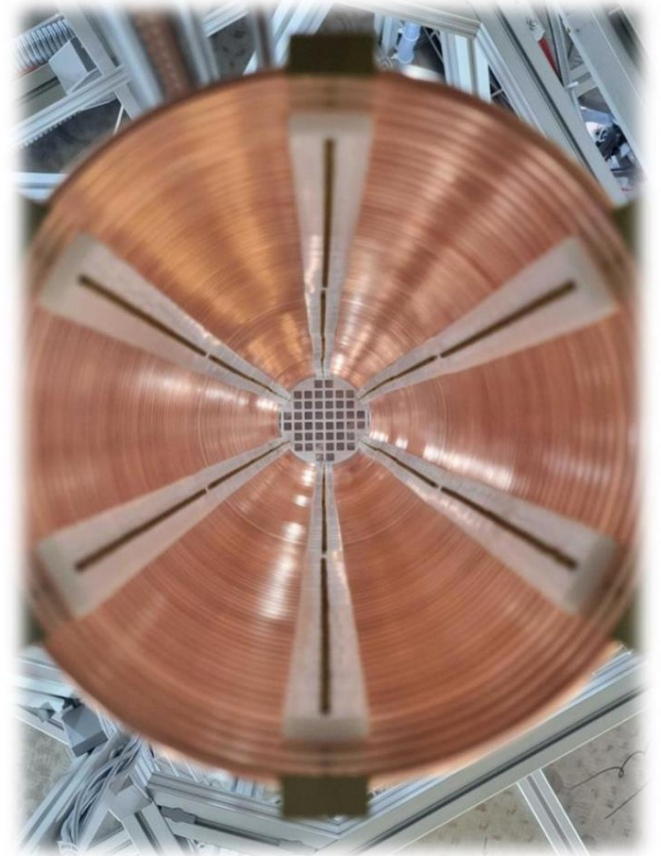
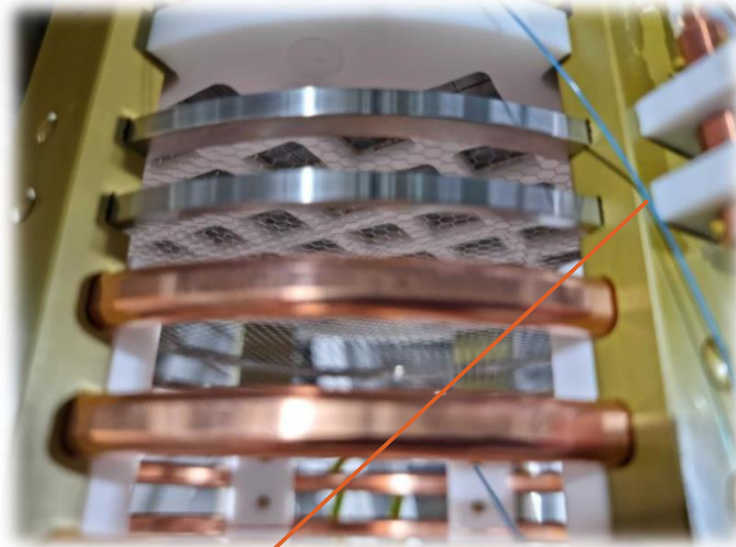
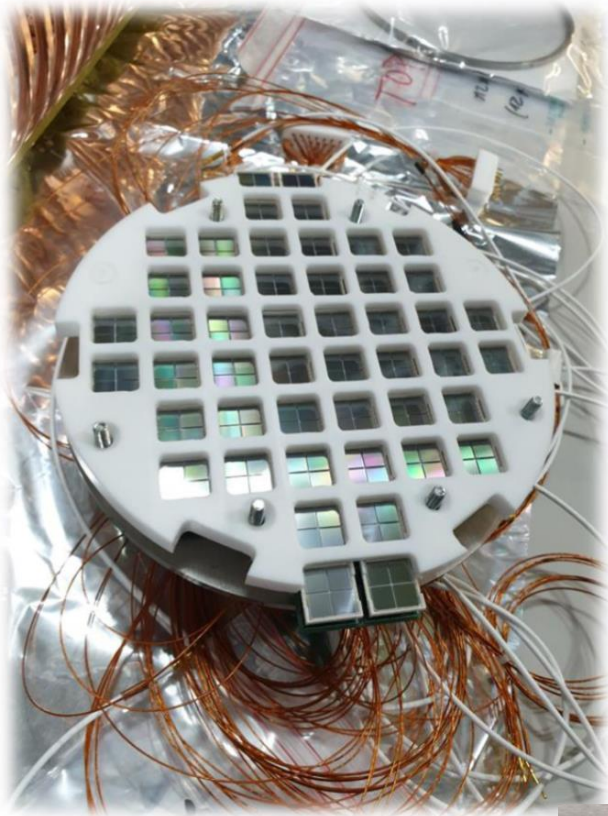


- Array = 12 tiles (or 192 SiPM cells) 12 x 12 mm² each
- Hamamatsu S13371-6050CQ-02 MPPC VUV SiPM
- Readout is made in parallel, so we have 12 channels
- 20x amplification per tile



[JINST 18 C03027 \(2023\)](#)

Installing the array in Xenoscope



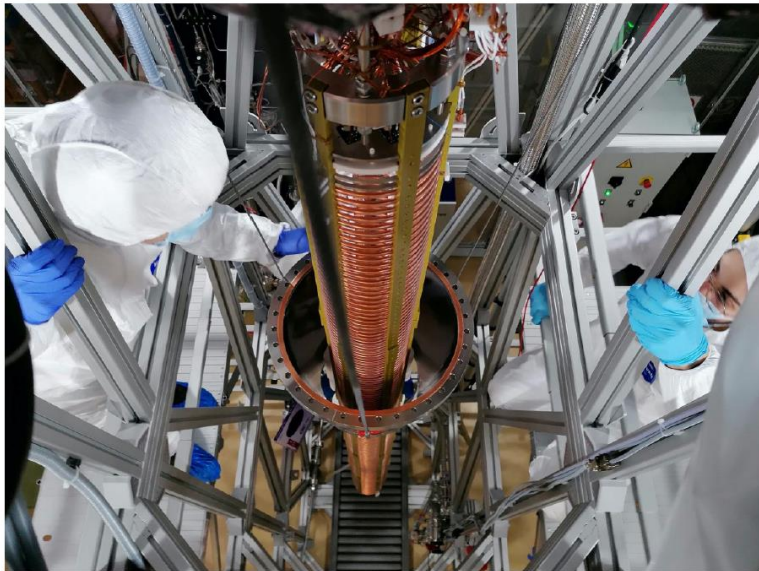
Calibration system:

- Blue light emitting diode (~460 nm) inside black box on air side
- Light transmitted via optical silica fibre

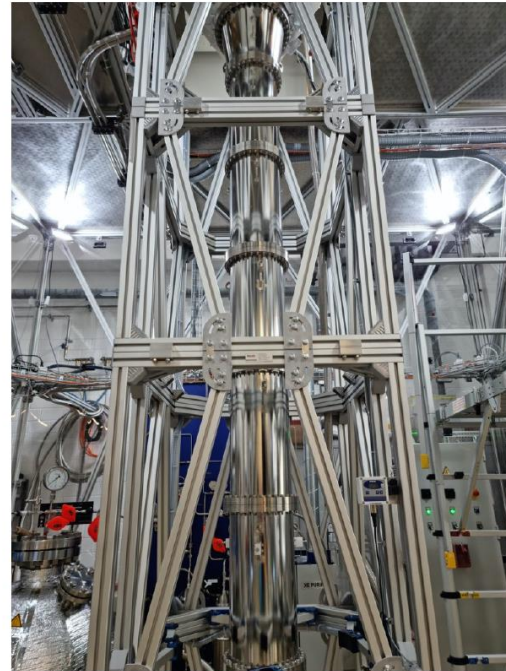


We have 2 fibres (180°)

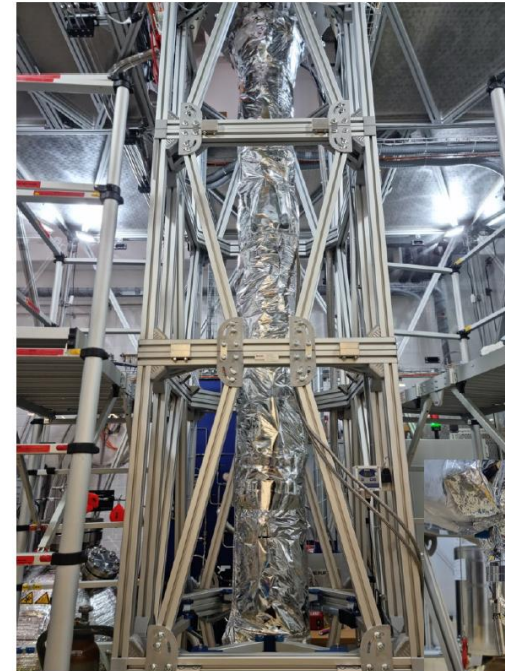
Putting all together



Connect all modules, cables, sensors...and close inner cryostat



Insulate with mylar and install the vacuum-to-LXe feedthrough



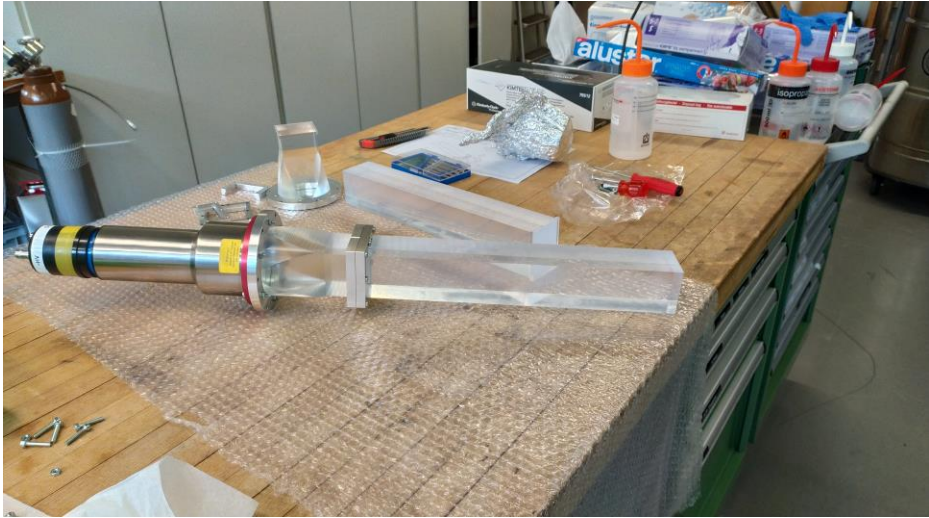
Close the outer cryostat and install high-voltage



Additional subsystems: 1. Muon detectors

For trigger in coincidence

Here they are



2. Liquid xenon recuperation: BoX (Ball of Xenon)

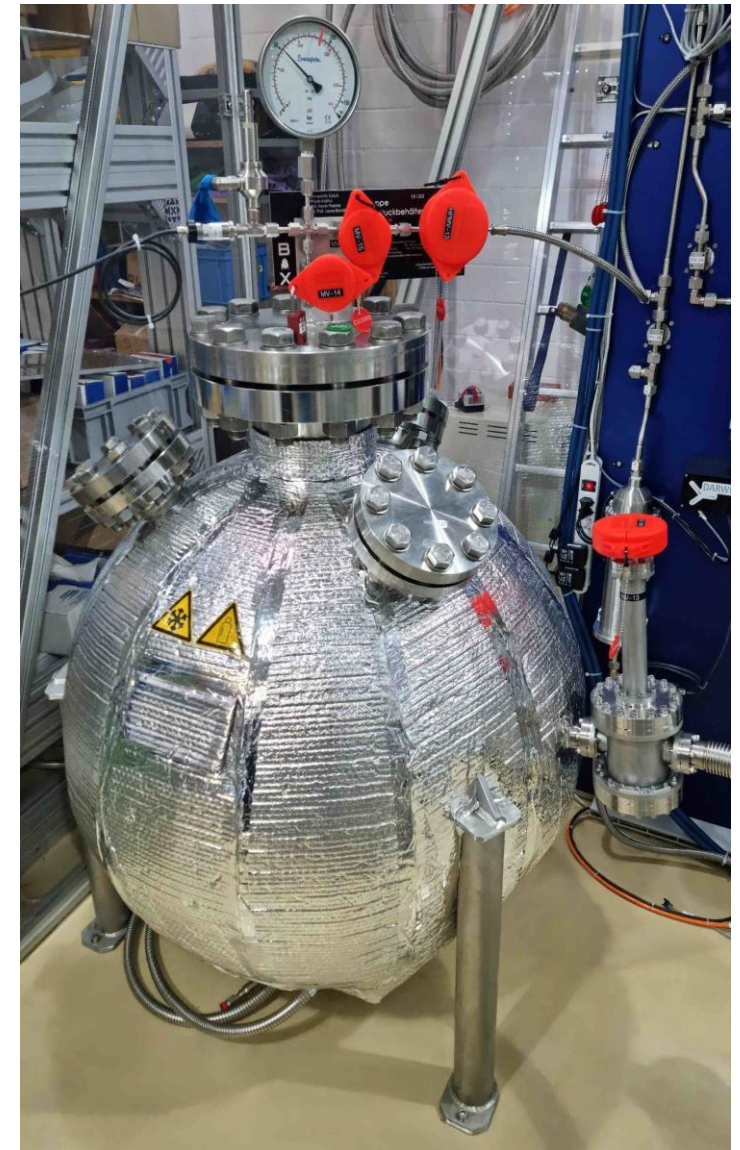
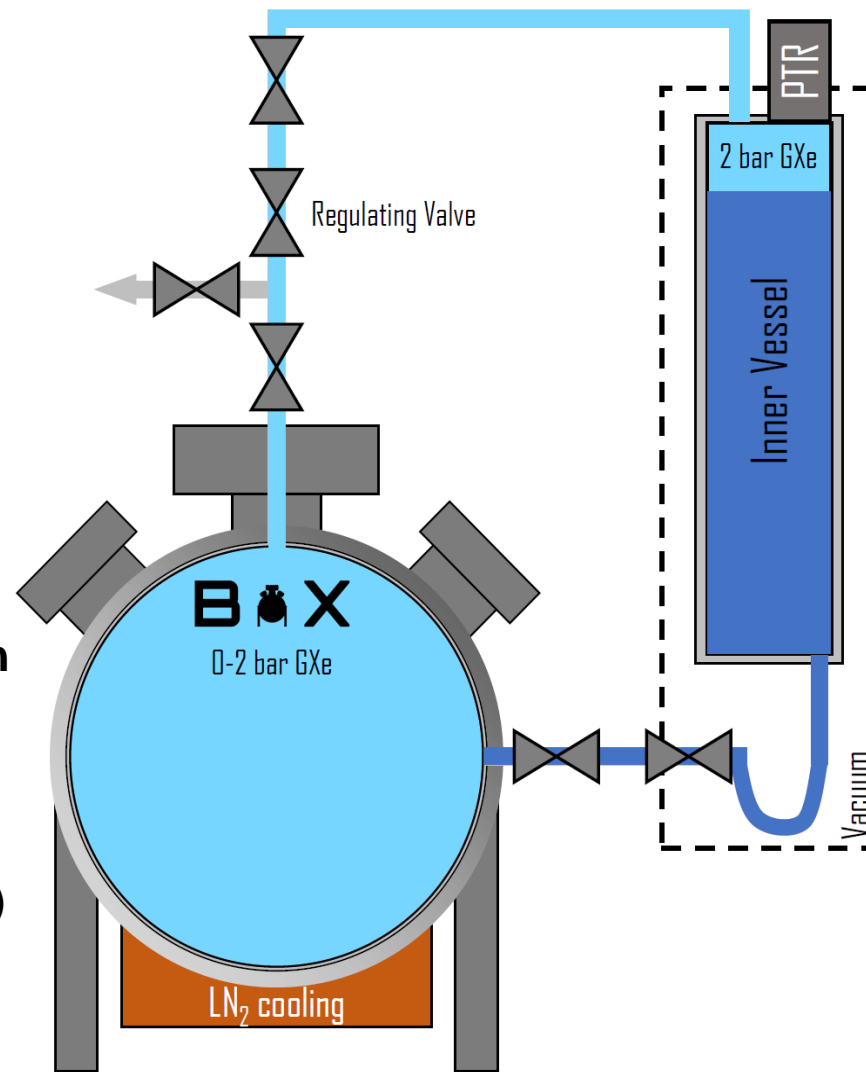
- Stainless Steel spherical vessel
- Gravity-assisted LXe recuperation system
- It can hold 450 kg of xenon at room temperature (max. pressure of 90 bar)

The recuperation speed is ~51 kg/h (only using gas phase ~19 kg/h)

For reference:

Using BoX the entire system (350 kg) was recuperated in 12 h (including LN₂ cooling process)

Without BoX it takes two-three days



3. Moving Xenon: the compressor

Double diaphragm KNF compressor:

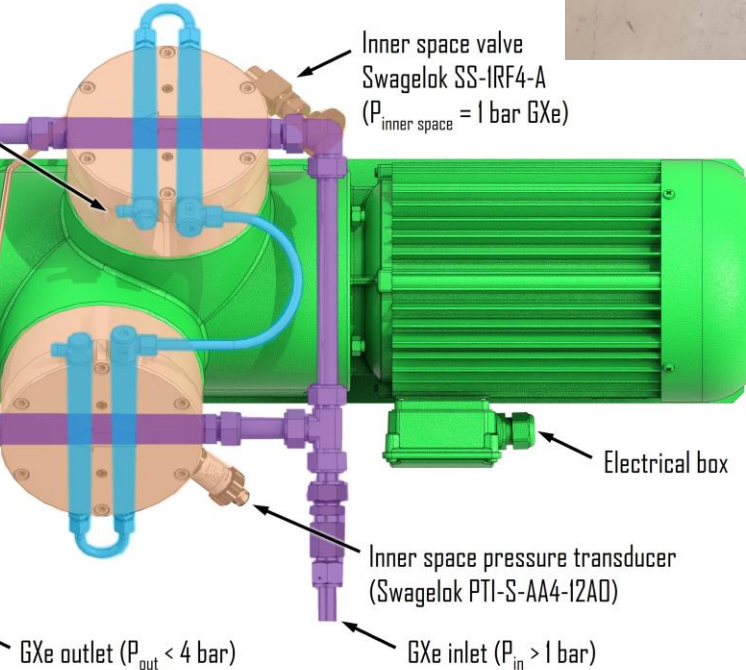
- Modified to operate at 100 slpm
- Should be operated with **at least** 1 bar abs. at the inlet
- Too much heat can be induced in the heads, causing the degradation of the diaphragm membranes

KNF PM32352-N1400.1.2.12E

Water cooling inlet and outlet are interchangeable

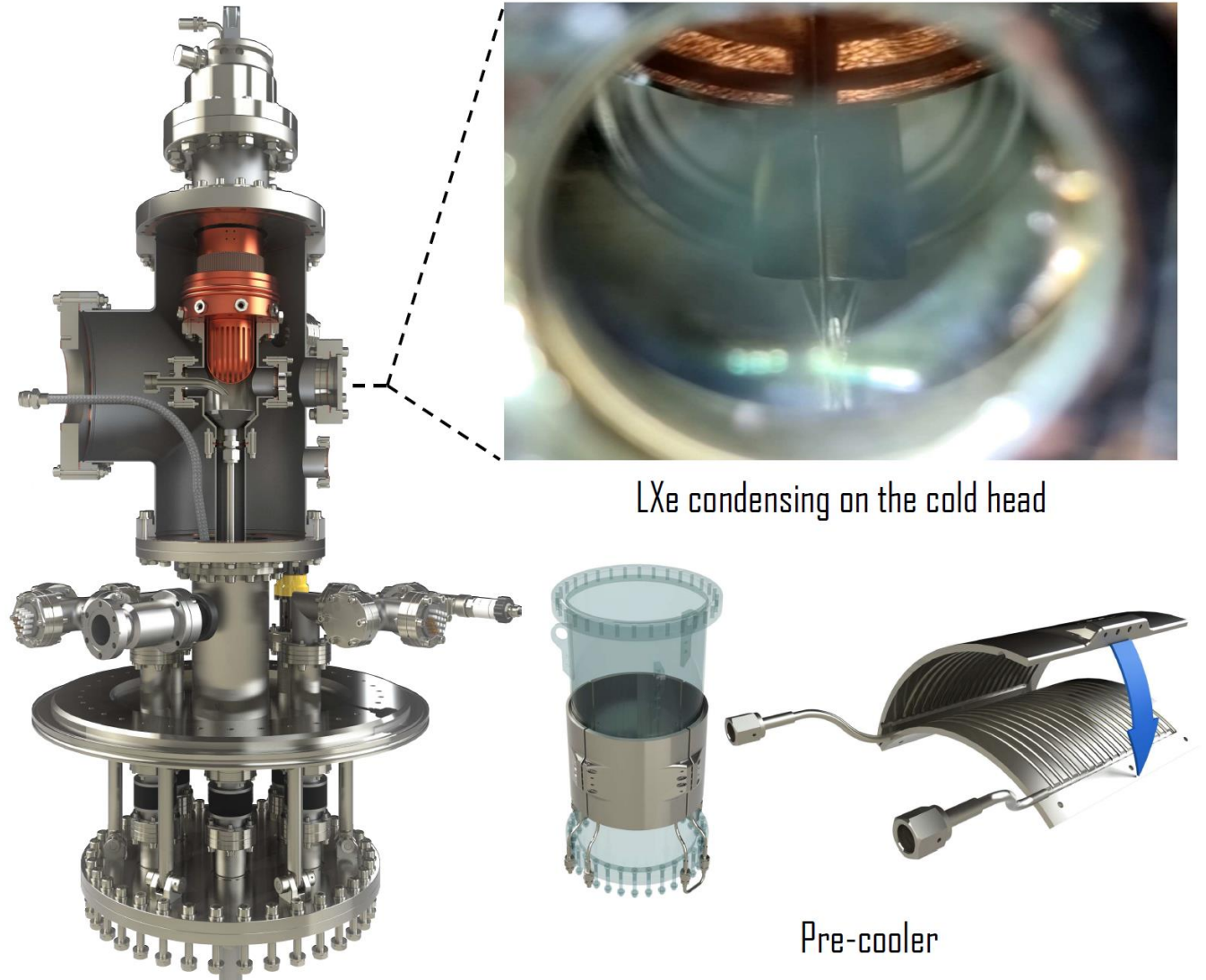
Inner space of both heads are connected together

Legend
Inner Space
Water Cooling
GXe

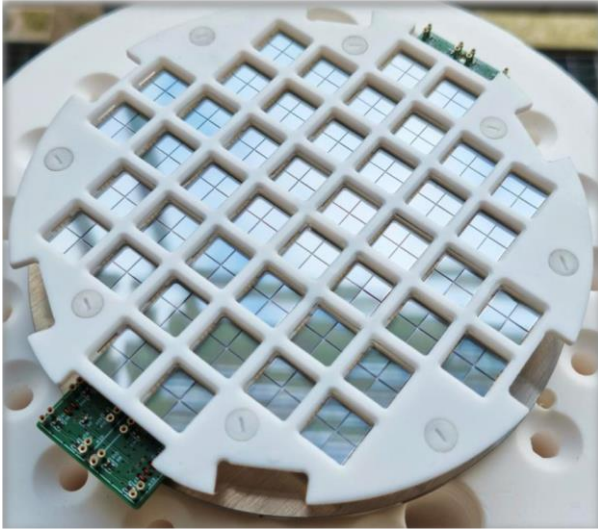


4. Cooling things

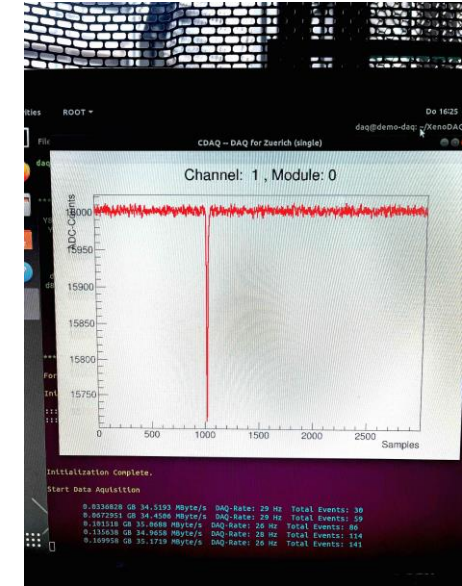
- Iwatani PC-150 PTR (~200 W cooling) and heater (~180 W heating)
- LN2 pre-cooler attached to the inner vessel for additional cooling power:
 - Filling w/o pre-cooler: 4 kg/h
 - Filling with pre-cooler: 17 kg/h
- Unpurified Xe extracted from the top of the inner vessel
- LXe condensing on the cold head
- Funnel carries purified Xe to the bottom of the inner vessel



5. Taking data



12 signal channels

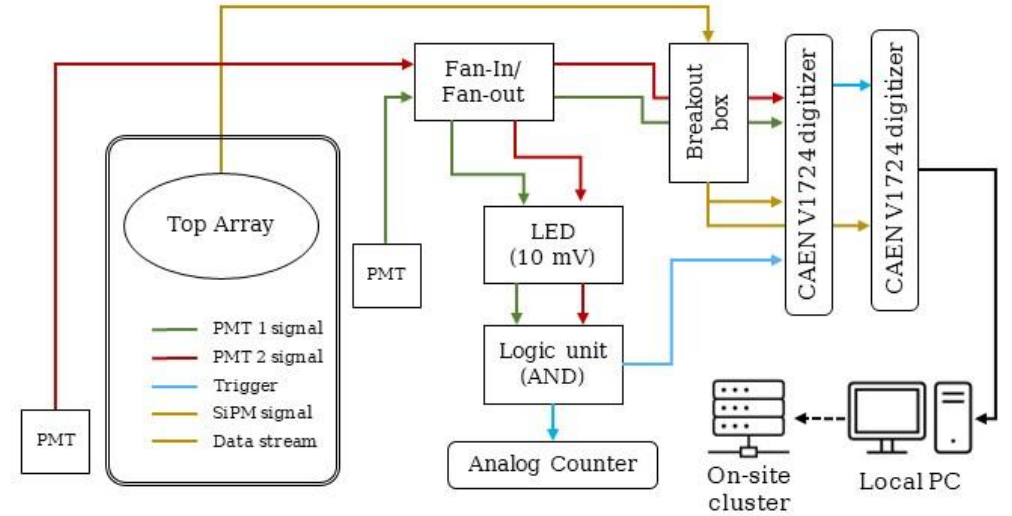
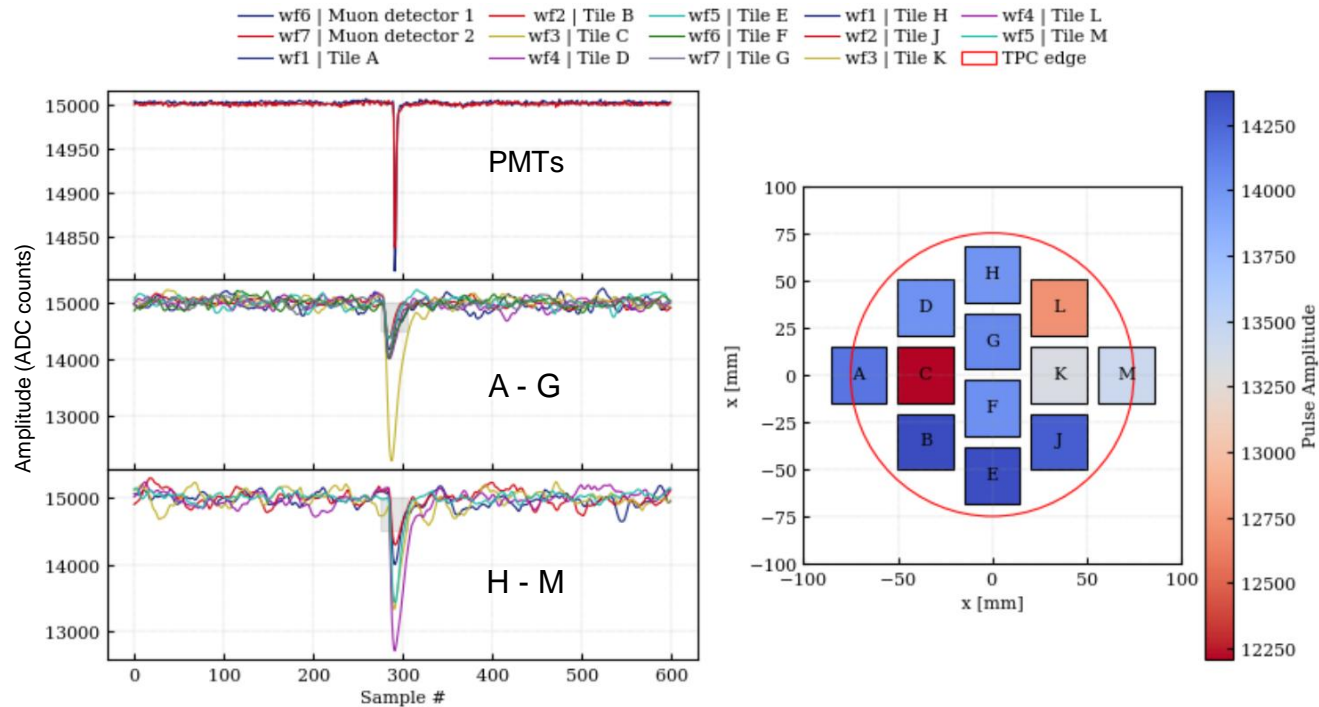


- CAEN V1724
- Two 8-channel boards daisy-chained
- 14 bits
- 100 MS/s
- 0-2.25 V Dynamic range

- DAQ interface configured through a XML file
- Output data in ROOT format
- Data sent to local cluster for storing

First test of the system: muon coincidence in gas

- The system was filled with 2 bar of GXe at room temperature
- Data acquired with a dual-coincidence trigger between the two PMTs coupled to the scintillators
- The total acquired livetime was 21 days



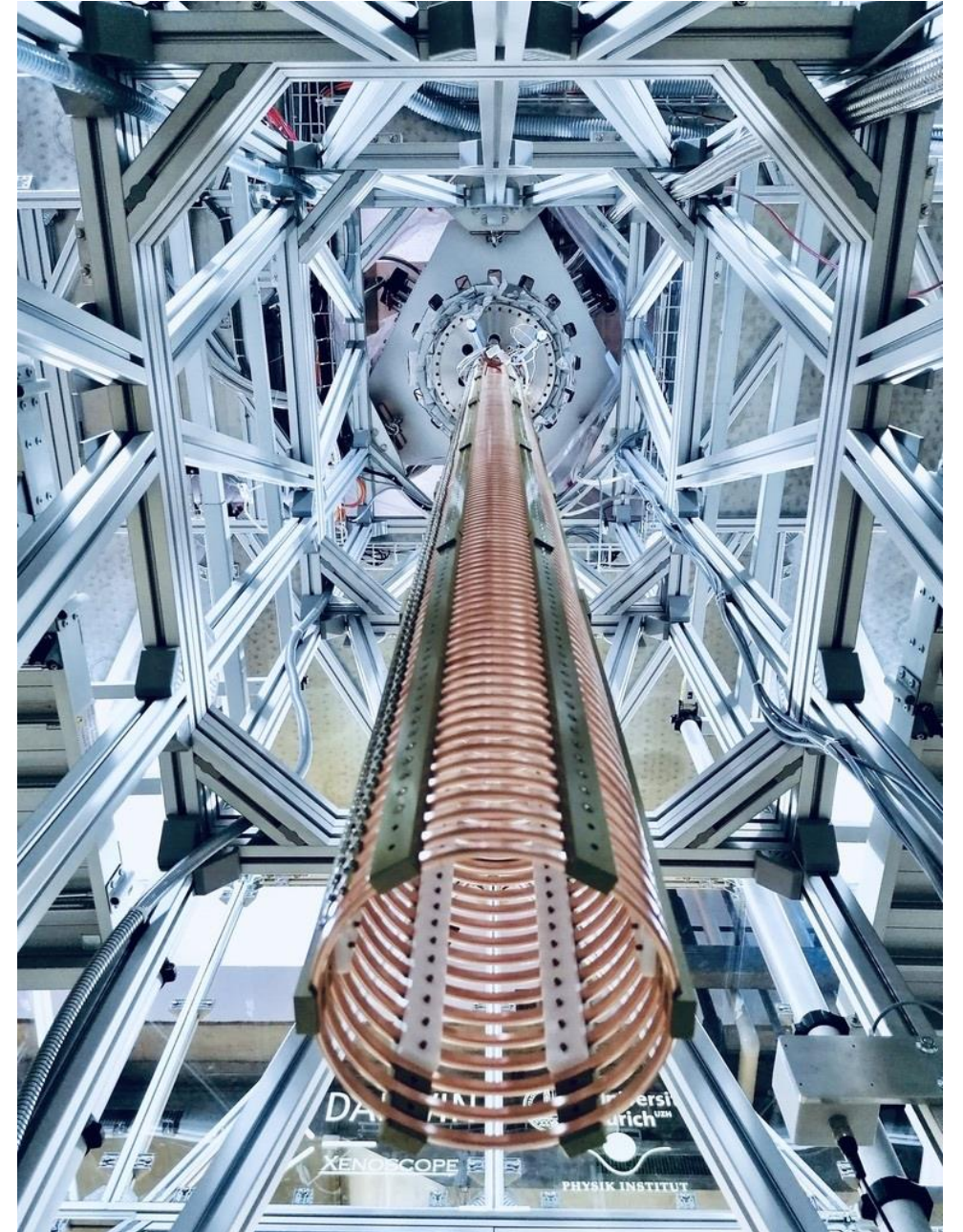
It works!

Summary

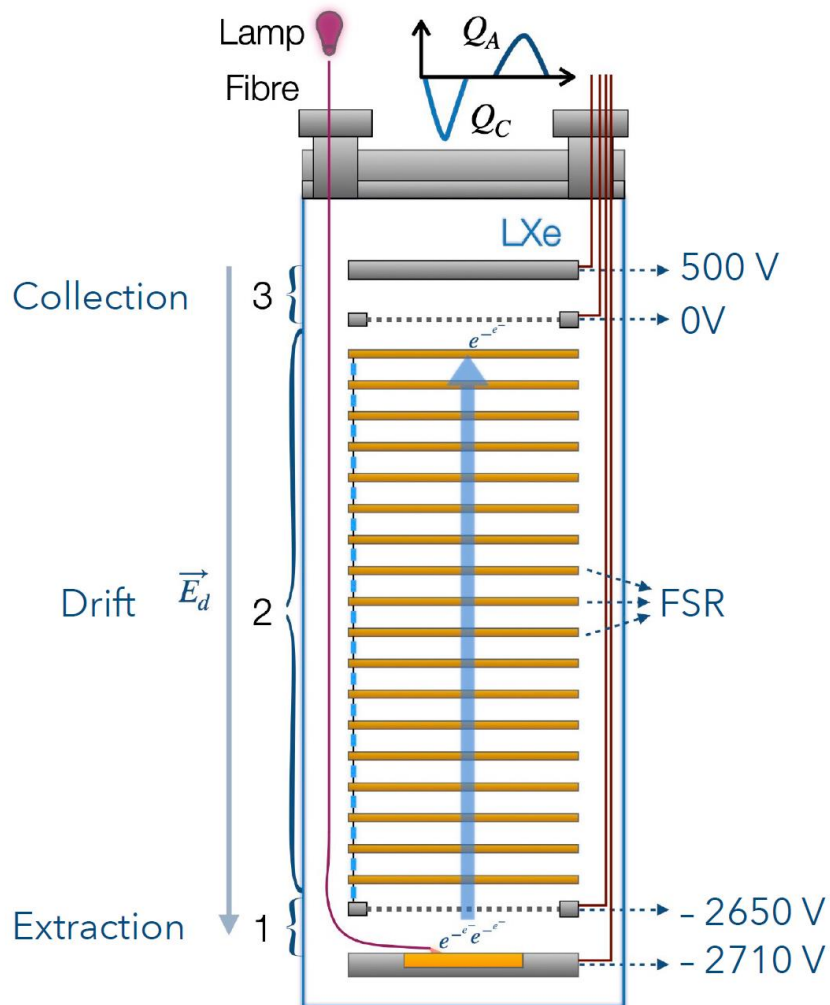
- We have built Xenoscope, a full-height DARWIN vertical demonstrator
- One first run (purity monitor) was successfully conducted
- Several upgrades were done for the full 2.6-m TPC
- The SiPM array was successfully tested in GXe
- The system is ready to be operated soon in liquid

The Xenoscope team:

Marta Babicz, Laura Baudis, Alexander Bismark, Jose Cuenca García, Paloma Cimental, Ricardo Peres, Mariana Rajado Silva, Diego Ramírez García, Christian Wittweg.

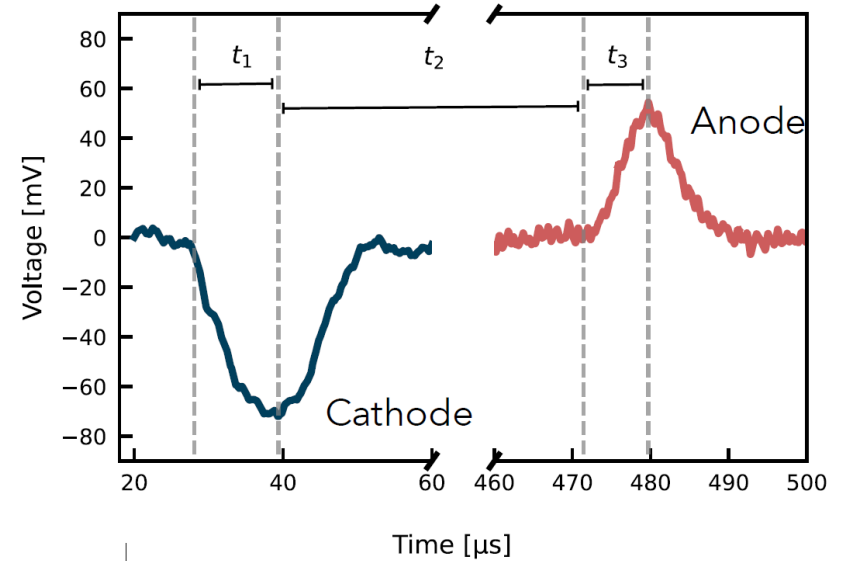


PM run: Form of the signals



- 343 kg of xenon
- Measurements at 2 bar and 177.6 K
- Illumination at 1 Hz
- First observable waveforms after 26.5 days of purification

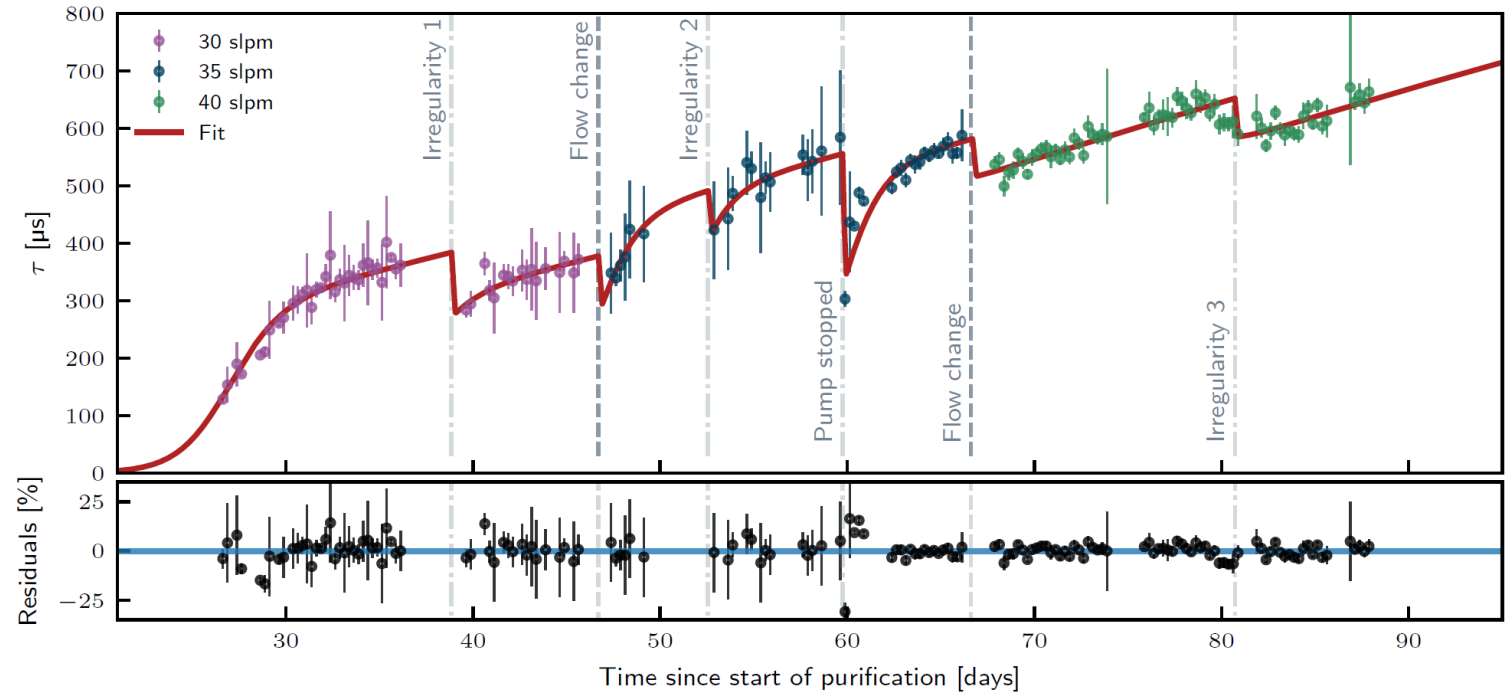
$$\frac{Q_A}{Q_C} = \frac{t_1}{t_3} e^{-(t_1+t_2+t_3)/\tau} \frac{(e^{t_3/\tau} - 1)}{(e^{-t_1/\tau} - 1)}$$



- Pulse delay: **drift velocity**
- Relative pulse area: **electron drifttime**
- Pulse shape: **longitudinal cloud diffusion**

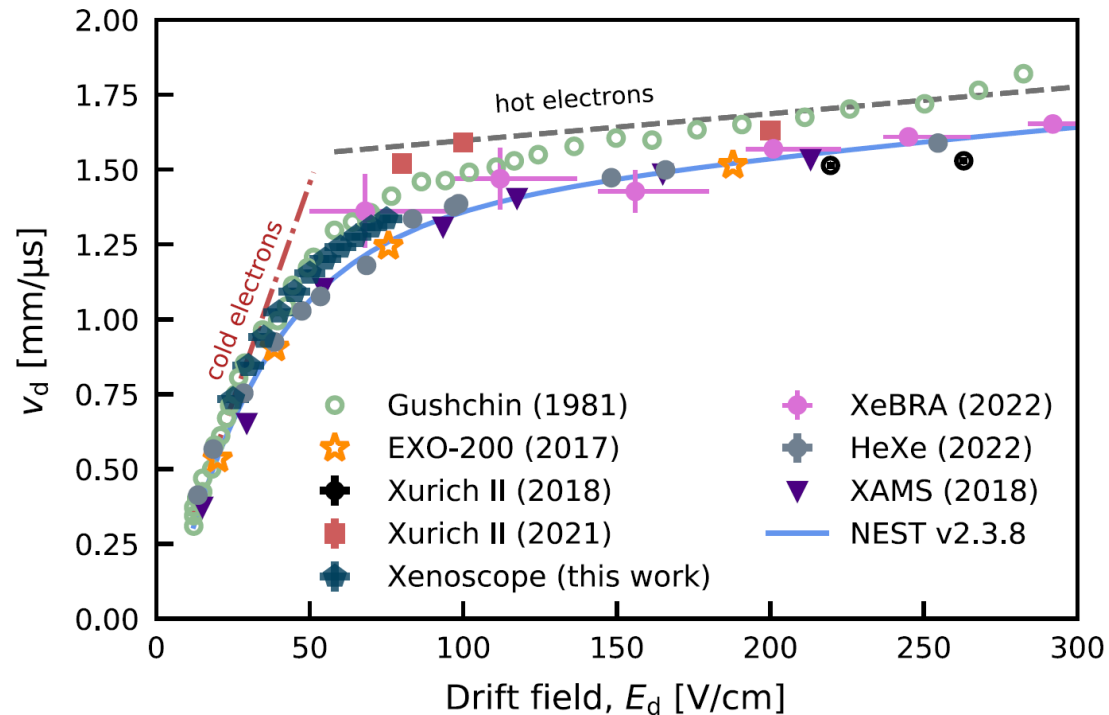
PM run: Purification performance

- Three different fluxes during the run:
 - 30 slpm (46.6 d)
 - 35 slpm (20.0 d)
 - 40 slpm (21.2 d)
- Electron drift lifetime drops during flow changes due to the liberation of trapped impurities at the LXe surface collar.
- Reached **(664 ± 23) μs** comparable to LUX and XENON1T
- Expected to reach > 2 ms with gas phase purification and 80 slpm



$$M_1 \frac{dI_1^{(j)}}{dt} = \underbrace{-F_1 \rho I_1}_{\text{Purification}} + \underbrace{\left(\frac{\Lambda_{1,0}}{1 + \frac{t - \Delta t_1^{(j)}}{T_{1/2,1}}} + C_1 \right)}_{\text{Outgassing}} - \underbrace{\frac{\varepsilon_1 P_C I_1}{h} + \frac{\varepsilon_2 P I_g}{h}}_{\text{Gas-liquid impurity exchange}} + \underbrace{M_1 \Delta I_1^{(j)} \int \delta(t - t^{(j)}) dt}_{\text{Discontinuities}}$$

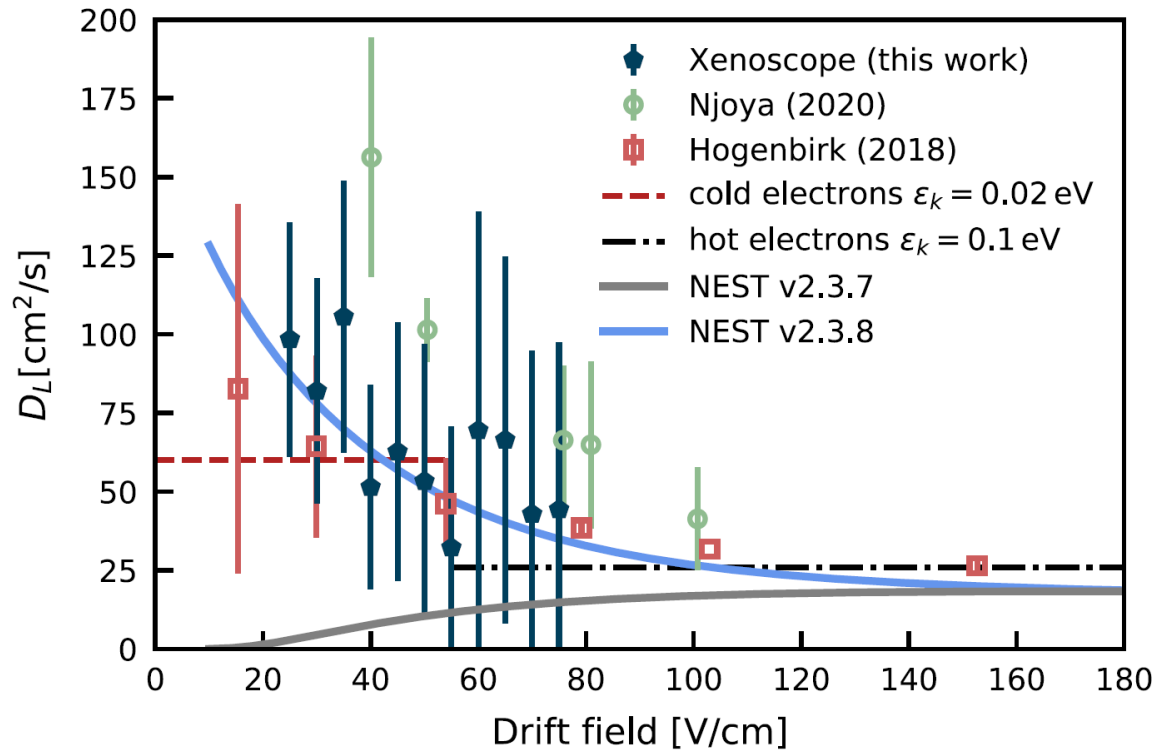
PM run: Electron drift velocity



$$v_d = \mu E_d = \frac{d}{t}$$

- Drift velocity from $Ed = 25 - 75$ V/cm
- Measurement at the end of the purification campaign at constant electron lifetime
- Electron mobility μ gives time between elastic collisions of electrons with xenon (and impurities, also inelastic)
- Depends on temperature, density and electron energy
- Impurities can offer more efficient energy loss than scattering on xenon leading to higher mobility

PM run: Longitudinal diffusion



Anode Gaussian peak width

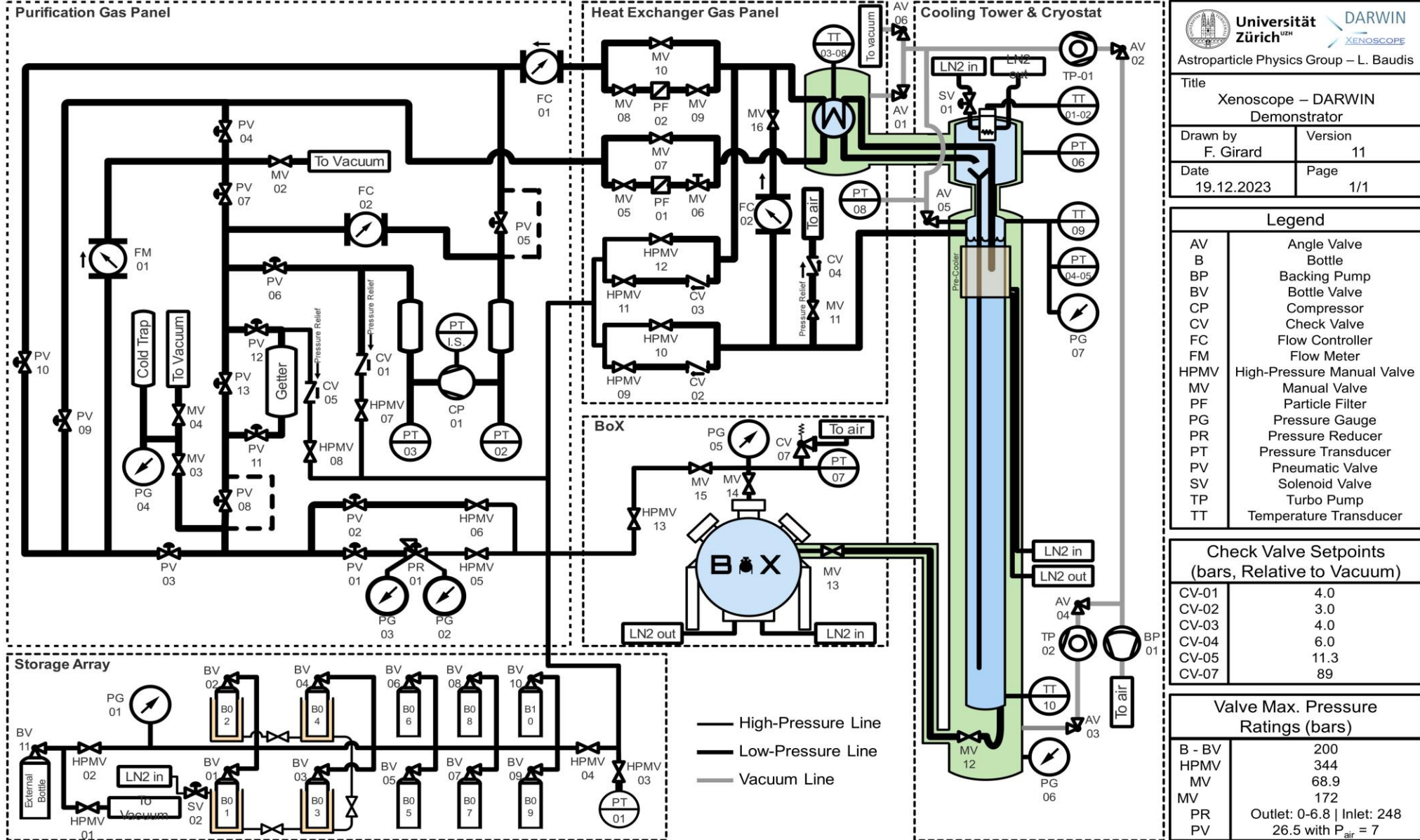
$$\sigma^2 = \frac{D_L 2t^3}{d^2} + \sigma_0^2$$

Diffusion coefficient, drift length, drift time

Cathode Gaussian peak width

$$D_L = \frac{(\epsilon_T + \epsilon)}{e} \mu$$

Systematic effect	Treatment and uncertainty
Measured	
Anode signal width, σ	Gaussian plus sine fit, 2 – 3%.
Drift time, $t_2 + t_3$	Time interval between extrema of the cathode and anode signal fits.
Initial signal width	Introduced by the lamp pulse. Measurements in vacuum and in LXe, (2.4 ± 0.2) μ s.
Electronics	RC time constant calculation from a square pulse, 0.2 μ s.
Drift length, $d_2 + d_3$	Drift distance of the electron cloud, taken as (513 ± 7) mm when accounting for the potential contraction of the components at 177 K with an assumed 1% thermal contraction, and the position of the centre of the cloud distribution in drift regions with respect the extrema of the signals.
Filtering and processing	Maximum 4% of anode signal width.
Simulated	
Detector response	COMSOL 3D model of the detector to derive the weighting potential, 10% uncertainty in the response.
Assumed	
Coulomb repulsion	Calculated with empirical model from [48], assumption of additional 5% uncertainty in the initial signal spread.
Electron attachment	Neglected, 4th-order correction: $\frac{D_L}{(v_d \cdot d)} \ll 1$.



Universität Zürich ^{uzh} DARWIN XENOSCOPE
 Astroparticle Physics Group – L. Baudis

Title: Xenoscope – DARWIN Demonstrator

Drawn by: F. Girard | Version: 11

Date: 19.12.2023 | Page: 1/1